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## Muon decay

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Polarisation and  $\gamma^5$ 

The radiative decay

The rare decay

Outlook





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Direct search Sensitive up to  $\mathcal{O}\left(10^3\,\text{GeV}\right)$ 

Indirect search Sensitive up to  $\mathcal{O}\left(10^{13}\,\text{GeV}\right)^1$ 

<sup>1</sup>Model dependent

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### Importance of background



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	Experimental	Theoretical (4-Fermi)	
Normal $\mu \rightarrow \nu \bar{\nu} e$	TWIST $^{1}$ $\mathcal{O}\left(10^{-4} ight)$	NLO (polarised, MC) [Arbuzov 2001] NNLO (unpolarised, analytic) [Anastasiou, Melnikov, and Petriello 2007]	
Radiative $\mu \rightarrow \nu \bar{\nu} e + \gamma$	$egin{array}{c} MEG \ \mathcal{O}\left(1\% ight) \end{array}$	NLO (polarised, MC) [Fael, Mercolli, and Passera 2015]	
Rare $\mu \rightarrow \nu \bar{\nu} e + e^+ e^-$	$Mu3e^2 \mathcal{O}\left(10\% ight)$	LO (polarised, MC)	

<sup>1</sup>Michel parameters <sup>2</sup>Proposed

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### **Open theoretical issues**

#### Use muon as toy process: clean QED

- $\gamma^5$ : Treatment in dim-reg?
- Regularisation scheme dependency
- Large logs, esp.  $\tau \rightarrow \nu \bar{\nu} e + \gamma$  (3.5 $\sigma$  deviation!) [Fael, Mercolli, and Passera 2015]
- No NLO for rare decay





### Polarisation and $\gamma^5$

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- Two equivalent ways of introducing polarisations dependence
  - "Closing the trace"  $u(p)\bar{u}(p) = (\not p + m)\frac{1+\gamma^5 \not s}{2}$
  - Massive spinor helicity formalism

$$u_{\pm}(p) = \left|\ell^{\pm}\right\rangle + \frac{m}{\left\langle\ell^{\pm}\right|n^{\mp}\right\rangle} \left|n^{\mp}\right\rangle$$
$$\left|k^{\pm}\right\rangle = \frac{1 \pm \gamma^{5}}{2} u(k)$$

 $n \ {\rm is} \ {\rm related} \ {\rm to} \ s \ {\rm [See \ Ellis' \ lecture]}$ 

• Both introduce a  $\gamma^5$ !



- $\gamma^5=\mathrm{i}\gamma^0\gamma^1\gamma^2\gamma^3$  is not well defined in d dimensions
- There are at two sources of γ<sup>5</sup>:

• The 4-Fermi vertex  $j^{\mu}_{V-A}(a,b)= \bar{\psi}_a\gamma^\mu(1-\gamma^5)\psi_b$ : [Berman and Sirlin 1962]

$$j^{\mu}_{V-A}(a,b) = \underbrace{\bar{\psi}_a \gamma^{\mu} \psi_b}_{j^{\mu}} - \bar{\psi}_a \gamma^{\mu} \psi'_b$$

 $\psi'_b = \gamma^5 \psi_b$  corresponds to an electron with  $m = -m_e$ .

• Polarisation: Spinor helicity formalism in FDH (external particles in d = 4)





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The radiative decay

Fully differential  $\operatorname{NLO}$  predictions for MEG @ PSI

• 4-Fermi interaction, fierzed at the Lagrangian

$$\mathcal{L} = \mathcal{L}_{\mathsf{QED}} + \frac{G_F}{\sqrt{2}} j_{V-A}(\mu, e) \cdot j_{V-A}(\nu_{\mu}, \nu_{e})$$





- Get amplitudes from GoSam [Cullen et al. 2014]
- FKS subtraction [Frixione, Kunszt, and Signer 1996]
- Custom phase spaces for increased stability and FKS
- (Almost) original VEGAS for integration [Lepage 1980]

## Branching ratio: Experimental comparison



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## Branching ratio: Experimental comparison



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#### Theorist's version of the MEG detector @ PSI





Fig.: Angular distribution:



Fig.: Angular distribution: MEG cuts on the electron loose  $4.10\,\%$  of the events



Fig.: Angular distribution: Polarised source  $\langle\cos\theta_c\rangle\approx-0.063<0$  corresponding to  $\langle\theta_c\rangle\approx93.62^\circ$ 



### Neutrino spectrum



Fig.: Neutrino spectrum:





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Fig.: Neutrino spectrum: Experimental resolution  $\approx 2 \text{ MeV} \Rightarrow \text{low energy}$  neutrino are important (below 5 MeV:  $7.2 \times 10^{-4}$ )





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The rare decay

 $\rm NLO$  branching ratios for Mu3e @ PSI

- $4_{\text{Born}} + 40_{1-\text{loop}} + 20_{\text{real}}$  diagrams up to pentagons
- A lot but not that many
- Use same approach (GoSam, FKS, VEGAS)
- Phase space more important than ever
- Mu3e cuts  $E_e > 10 \text{ MeV}$





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# **Preliminary results**

	LO	NLO only	K-factor
$\mathcal{B}(no  cuts)$	$3.605 imes10^{-5}$	$0.007 imes10^{-5}$	-0.19%
$\mathcal{B}(E>10\mathrm{MeV})$	$2.309 imes10^{-6}$	-0.041 $ imes$ 10 <sup>-6</sup> $ $	-1.78%





### Polarisation and $\gamma^5$

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## Conclusion and outlook

Conclusion

- New fully differential  $\operatorname{NLO}$  predictions for the radiative decay for  $\mathsf{MEG}$
- $\bullet~{\rm New~NLO}$  predictions for the rare decay for Mu3e



Conclusion

- New fully differential  $\operatorname{NLO}$  predictions for the radiative decay for  $\mathsf{MEG}$
- New  $\operatorname{NLO}$  predictions for the rare decay for Mu3e
- Work to be done
  - Predict / Resum large logs  $\log \frac{m}{M} \log \frac{\omega_0}{M}$  @ NLO and possibly NNLO
  - $\Rightarrow$  Solve  $3.5\sigma$  discrepancy
    - Produce distributions for the rare decay @ NLO



## A slide from Andrea Visconti

Regularization scheme dependence of two-loop amplitudes

└\_<sub>Schemes</sub>

- **1** Variants of dimensional regularisation:
  - CDR ("conventional dimensional regularization")
  - HV ("'t Hooft Veltman")
- 2 Variants of dimensional reduction:
  - DRED ("original/old dimensional reduction")
  - FDH ("four-dimensional helicity scheme")

	CDR	$_{\rm HV}$	FDH	DRED
internal gluon	$\hat{g}^{\mu\nu}$	$\hat{g}^{\mu u}$	$g^{\mu\nu}$	$g^{\mu\nu}$
external gluon	$\hat{g}^{\mu u}$	$\bar{g}^{\mu u}$	$\bar{g}^{\mu u}$	$g^{\mu u}$
$\hat{g}_{a}$	$\hat{g}_{\hat{q}}$		g ¢ q	$\int_{a}^{g} q$
ĝ for ĝ	6000	<u> </u>	600	g for

FDH

g=Q4D ,  $\hat{g}=QD-{\rm dim.}$  ,  $\bar{g}=4-{\rm dim.}$ 

HV

CDR

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DRED